## Dark Matter - Analyzing Galaxy UGC 11748

We can see Galaxy UGC 11748. Most of the stars lie within a radius of $1.64 \times 10^{20} \mathrm{~m}$. Physicists have been able to calculate the mass of these stars by the Brightness Method. The mass has been determined to be $1.54 \times 10^{41} \mathrm{~kg}$ (or 77.4 billion times the mass of our sun!!). We call this the 'luminous mass'. Why?


It is expected that the stars that lie outside this radius will orbit in the same way that planets orbit the sun. That is, the Fc needed for objects to follow a circular path is supplied completely by Fg . ( $\mathrm{Fg}=\mathrm{mg}$ does NOT work out in space - it is a short cut for objects on earth only).

The focus of this activity is to analyze these far stars.

## Diagram

## Useful formulas

Fnet $=m a \quad F c=m v^{2} / r$
$\mathrm{Fg}=\mathrm{Gm}_{1} \mathrm{~m}_{2} / \mathrm{r}^{2}$
Data

| Star of interest |  |  |  |  |
| :---: | :---: | :---: | :---: | :---: |
| Orbital radius <br> $\left(\times 10^{20} \mathrm{~m}\right)$ | Measured speed <br> $\left(\times 10^{5} \mathrm{~m} / \mathrm{s}\right)$ | Calculated speed <br> $\left(\times 10^{5} \mathrm{~m} / \mathrm{s}\right)$ | Gravitational Mass <br> $(\mathrm{kg})$ | Missing Mass <br> $(\%)$ |
| A: 1.85 | 2.47 | 2.36 | 1.69 | 8.99 |
| B: 2.75 | 2.40 |  |  |  |
| C: 3.18 | 2.37 |  |  |  |
| D: 4.26 | 2.25 |  |  |  |
| E: 6.48 | 2.47 |  |  |  |

1. Graph - Use the data above to graph measured speed (y axis) vs. orbital radius ( $x$ axis). The graph should be at least $1 / 2$ page big. The $y$ axis should go from 0 to 2.5 with the unit being ( $x 10^{5} \mathrm{~m} / \mathrm{s}$ ). The $x$ axis should go from 0 to 7 with the unit being ( $\times 10^{20} \mathrm{~m}$ ). Make sure you label the axes.

Draw a line or curve of best fit (refer to page $\qquad$ In your text) . Label this 'measured'.
2. For each orbital radius (see chart), calculate the speed expected if the only mass causing this rotation is the luminous mass. Record your answers in the column 'Calculated speed'. One answer is provided for you. You should do the math and make sure you get the same answer! If you do, you know your math is good.
a) Show a sample calculation
b) Plot the calculated speed against orbital radius on the graph you've already started (\#1). Draw a line/curve of best fit. Label this 'calculated'.
3. Compare the 'measured' and 'calculated' lines on the graph. Discuss a possible explanation for any difference.
4. Gravitational Mass: The measured speeds are accurately measured. Let's work backwards and determine what mass is responsible for the force of gravity pulling these stars in a circle. Set Fc = Fg and think about what $m_{1}$ and $m_{2}$ represent. It will be helpful to go back to $\mathrm{Fc}=\mathrm{mv}^{2} / \mathrm{r}$ calculations and see what the variables represent.
a) Show a sample calculation.
b) Record your answers in the chart in the 'Gravitational Mass' column.
5. For each star, calculate the \% difference between the gravitation mass and the luminous mass.

Represent this difference as a percentage of the gravitational mass.
a) Show a sample calculation.
b) Record you answers in the 'missing mass' column.
6. Explain the significance of the missing mass.

Application - 15 marks.

Name:
$\square$

| Application |  |
| :--- | :--- |
| \#1, \#2 Graph |  |
| \#3 - 3 marks |  |
| \#4 - 3 marks |  |
| \#5 - 2 marks |  |
| \#6-2 marks |  |

